

# Magnetic field extrapolation in close binary stars

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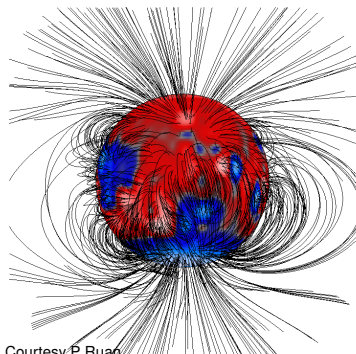
Kiepenheuer-Institut für Sonnenphysik, Freiburg i.Br.

Binamics meeting, Vienna, Feb 2016

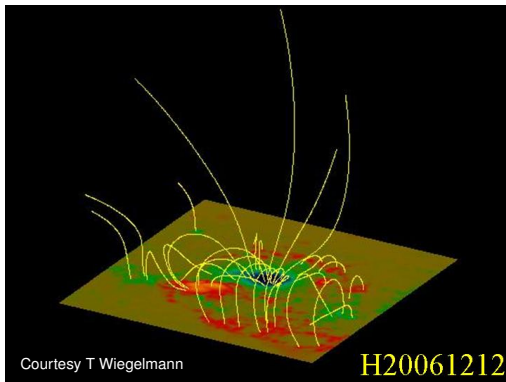
# Solar magnetic fields

## Coronal field extrapolation

- observation of coronal magnetic field hardly possible
- ⇒ approximation techniques:
- **current-free** (potential) field approx.
  - **non-/linear force-free** field approx.
  - **magneto-hydrostatic** stratification



Courtesy P Ruan

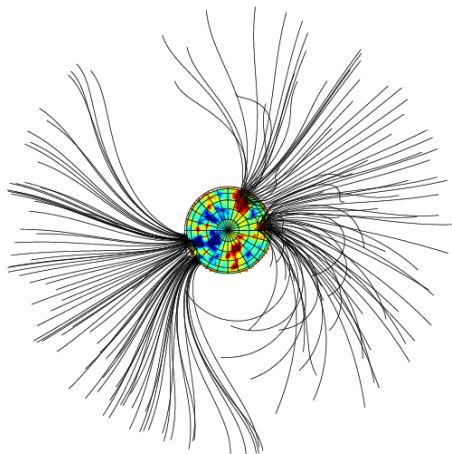
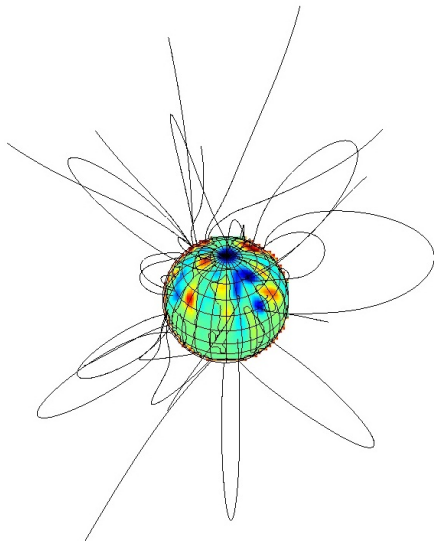


Courtesy T Wiegelmann

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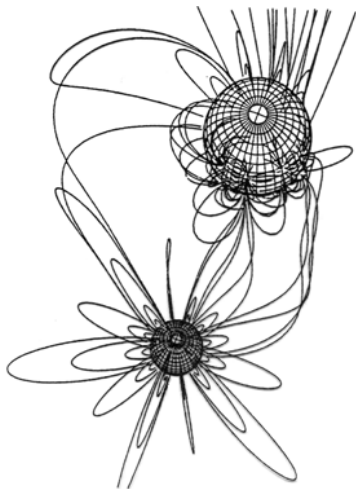
# AB Dor

Potential field source surface (PFSS) approximation of coronal field structure (Jardine et al. 2002)



# Joint magnetospheres of close star systems

## Project overview



Uchida & Sakurai 1985

- targets:
  - e.g.: RS CVn-/BY Dra-systems
  - star-exoplanet systems
  - circumbinary exoplanets
  - hot binary stars
  - ...
- objectives:
  - structure of joint magnetosphere
  - open/closed/inter-connecting magnetic flux
  - magnetic interactions
  - atmospheric structure
  - observational signatures
  - ...

# Theoretical approach

## PFSS approximation

- Assumption: **current-free** magnetic field

$$\nabla \times \mathbf{B} = 0 \quad \rightarrow \quad \mathbf{B} = -\nabla\psi \quad \psi: \text{scalar flux function}$$

- Solenoidal field:

$$\nabla \cdot \mathbf{B} = 0 \quad \rightarrow \quad \Delta\psi = 0$$

→ **Laplace equation**

# Theoretical approach

## PFSS approximation

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$$\nabla \cdot \mathbf{B} = 0 \quad \rightarrow \quad \Delta\psi = 0$$

### → Laplace equation

- General solution:

$$\psi(\mathbf{r}) = \sum_{l,m} a_{lm} R_{lm}(\mathbf{r}) + b_{lm} I_{lm}(\mathbf{r})$$

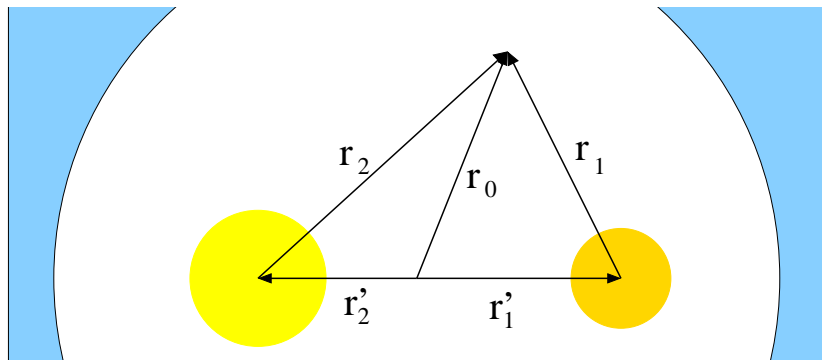
$a_{lm}, b_{lm}$  expansion coefficients;  $Y_{lm}$  spherical harmonics

$R_{lm}(\mathbf{r}) = r^l Y_{lm}(\theta, \phi)$       **regular** solid spherical harmonics

$I_{lm}(\mathbf{r}) = \frac{1}{r^{l+1}} Y_{lm}(\theta, \phi)$       **irregular** solid spherical harmonics

# Application to binary systems

## Superposition of flux functions



$$\Psi(\mathbf{r}) = \Psi^{(0)}(\mathbf{r}_0) + \Psi^{(1)}(\mathbf{r}_1) + \Psi^{(2)}(\mathbf{r}_2)$$

- $\Psi^{(1)}(\mathbf{r}_1) = \mathbf{C}^{(1)} \cdot \mathbf{l}(\mathbf{r}_1)$  'sources' inside **primary**
- $\Psi^{(2)}(\mathbf{r}_2) = \mathbf{C}^{(2)} \cdot \mathbf{l}(\mathbf{r}_2)$  'sources' inside **secondary**
- $\Psi^{(0)}(\mathbf{r}_0) = \mathbf{C}^{(0)} \cdot \mathbf{R}(\mathbf{r}_0)$  'mirror sources' outside **source surface**

# Application to binary systems

## Boundary conditions

- *Boundary conditions:*
  - **Source surface:** Stellar winds drag magnetic field in radial direction

$$\Psi \Big|_{S_0} = \Psi^{(0)} \Big|_{S_0}(\mathbf{r}_0) + \Psi^{(1)} \Big|_{S_0}(\mathbf{r}_1) + \Psi^{(2)} \Big|_{S_0}(\mathbf{r}_2) = \text{const.}$$

- **Primary** and **secondary:** observed radial magnetic field maps

$$B_r^{(1)} = -\frac{\partial}{\partial r_1} (\Psi^{(0)}(\mathbf{r}_0) + \Psi^{(1)}(\mathbf{r}_1) + \Psi^{(2)}(\mathbf{r}_2)) \Big|_{S_1}$$
$$B_r^{(2)} = -\frac{\partial}{\partial r_2} (\Psi^{(0)}(\mathbf{r}_0) + \Psi^{(1)}(\mathbf{r}_1) + \Psi^{(2)}(\mathbf{r}_2)) \Big|_{S_2}$$



# Application to binary systems

## Boundary conditions

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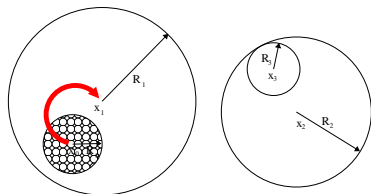
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- **Primary** and **secondary:** observed radial magnetic field maps

$$B_r^{(1)} = -\frac{\partial}{\partial r_1} \left( \Psi^{(0)}(\mathbf{r}_0) + \Psi^{(1)}(\mathbf{r}_1) + \Psi^{(2)}(\mathbf{r}_2) \right) \Big|_{S_1}$$
$$B_r^{(2)} = -\frac{\partial}{\partial r_2} \left( \Psi^{(0)}(\mathbf{r}_0) + \Psi^{(1)}(\mathbf{r}_1) + \Psi^{(2)}(\mathbf{r}_2) \right) \Big|_{S_2}$$

- *Problem:* Express  $\Psi^{(p)}$  in reference frame of sphere  $q$
- *Solution:* **Multipole translation theorems**

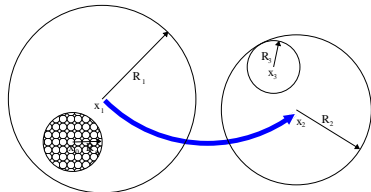
# Multipole translation theorems



**Irregular-irregular expansion:**  $|\mathbf{x} - \mathbf{x}_1| > |\mathbf{x}_1 - \mathbf{x}_0|$

$$I_{lm}(\mathbf{x} - \mathbf{x}_0) = \sum_{l'm'} (|I|)_{lm}^{l'm'}(\mathbf{x}_1 - \mathbf{x}_0) I_{l'm'}(\mathbf{x} - \mathbf{x}_1)$$

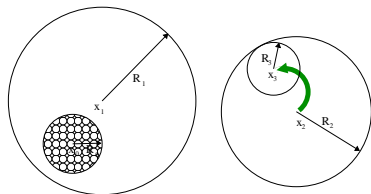
$$(|I|)_{lm}^{l'm'}(\mathbf{x}_1 - \mathbf{x}_0) = (-1)^{l'+m'-l-m} R_{l'-l}^{m-m'}(\mathbf{x}_1 - \mathbf{x}_0)$$



**Irregular-regular expansion:**  $|\mathbf{x} - \mathbf{x}_2| < |\mathbf{x}_2 - \mathbf{x}_1|$

$$I_{lm}(\mathbf{x} - \mathbf{x}_1) = \sum_{l'm'} (|IR|)_{lm}^{l'm'}(\mathbf{x}_2 - \mathbf{x}_1) R_{l'm'}(\mathbf{x} - \mathbf{x}_2)$$

$$(|IR|)_{lm}^{l'm'}(\mathbf{x}_2 - \mathbf{x}_1) = (-1)^{l'-m'} I_{l+l'}^{m-m'}$$



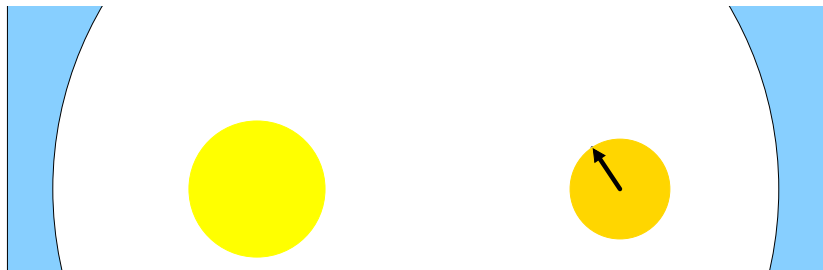
**Regular-regular expansion:**

$$R_{lm}(\mathbf{x} - \mathbf{x}_2) = \sum_{l'm'} (R|R)_{lm}^{l'm'}(\mathbf{x}_3 - \mathbf{x}_2) R_{l'm'}(\mathbf{x} - \mathbf{x}_3)$$

$$(R|R)_{lm}^{l'm'}(\mathbf{x}_3 - \mathbf{x}_2) = R_{l-l'}^{m-m'}$$

# Application to binary systems

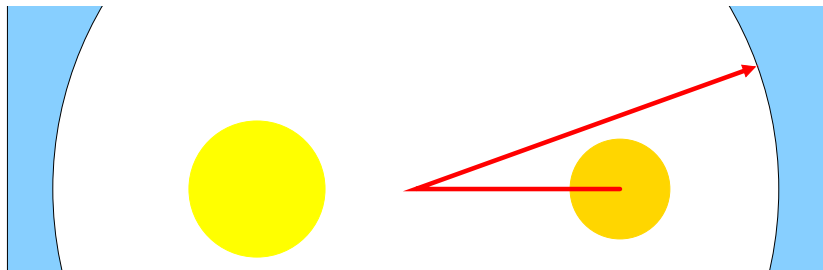
Example: Primary magnetic field



$$\mathbf{B}^{(1)} = -\mathbf{C}^{(1)} \cdot \left[ \mathcal{I}'_{a_1} - (||I)^{(10)} \cdot \mathcal{I}_{a_0} \cdot \mathcal{R}_{a_0}^{-1} \cdot (R|R)^{(01)} \cdot \mathcal{R}'_{a_1} \right] \\ - \mathbf{C}^{(2)} \cdot \left[ (||R)^{(21)} \cdot \mathcal{R}'_{a_1} - (||I)^{(20)} \cdot \mathcal{I}_{a_0} \cdot \mathcal{R}_{a_0}^{-1} \cdot (R|R)^{(01)} \cdot \mathcal{R}'_{a_1} \right]$$

# Application to binary systems

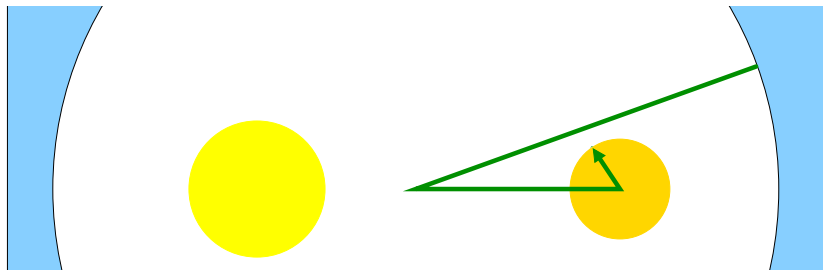
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# Application to binary systems

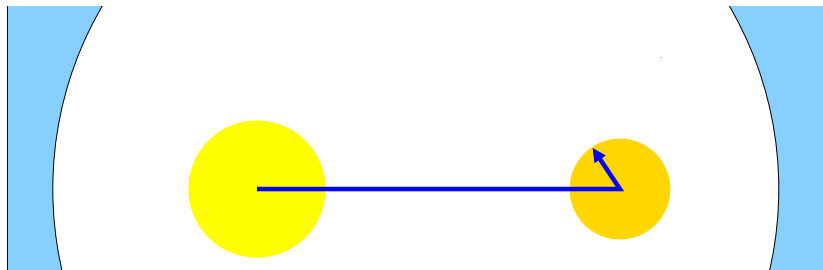
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# Application to binary systems

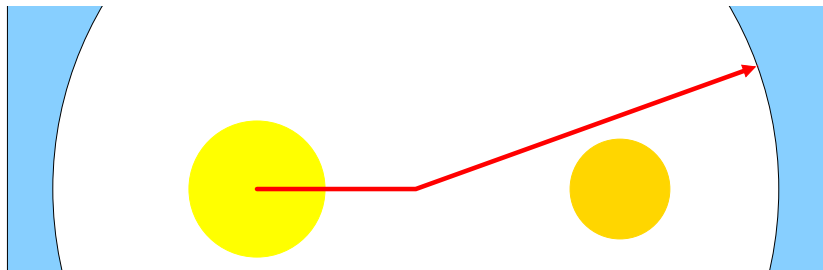
Example: Primary magnetic field



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# Application to binary systems

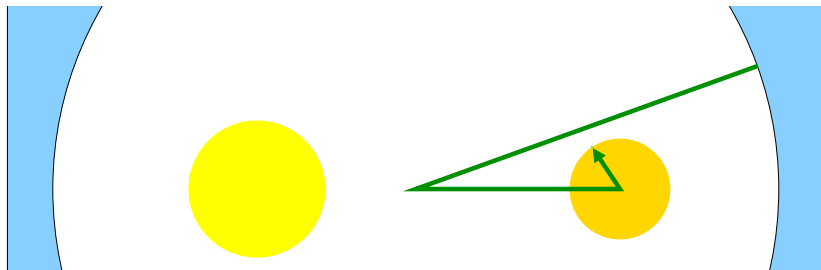
Example: Primary magnetic field



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# Application to binary systems

Example: Primary magnetic field



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# Application to binary systems

## Linear system of equations

- binary system:

$$\begin{pmatrix} \mathbf{B}^{(1)} \\ \mathbf{B}^{(2)} \end{pmatrix} = \begin{pmatrix} \mathcal{D}^{(11)} & \mathcal{D}^{(12)} \\ \mathcal{D}^{(21)} & \mathcal{D}^{(22)} \end{pmatrix} \cdot \begin{pmatrix} \mathbf{C}^{(1)} \\ \mathbf{C}^{(2)} \end{pmatrix}$$

- magnetic maps  $\mathbf{B}^{(1/2)}$  known  $\rightarrow$  solve for coefficients  $\mathbf{C}^{(1/2)}$

# Application to binary systems

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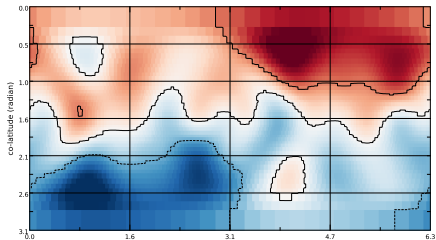
- magnetic maps  $\mathbf{B}^{(1/2)}$  known  $\rightarrow$  solve for coefficients  $\mathbf{C}^{(1/2)}$
- *post scriptum*: technique easily expandable to  $N$  objects

$$\begin{pmatrix} \mathbf{B}^{(1)} \\ \vdots \\ \mathbf{B}^{(j)} \\ \vdots \\ \mathbf{B}^{(N)} \end{pmatrix} = \begin{pmatrix} \mathcal{D}^{(11)} & \dots & \mathcal{D}^{(1j)} & \dots & \mathcal{D}^{(1N)} \\ \vdots & \ddots & \vdots & & \vdots \\ \mathcal{D}^{(j1)} & \dots & \mathcal{D}^{(jj)} & \dots & \mathcal{D}^{(jN)} \\ \vdots & & \vdots & \ddots & \vdots \\ \mathcal{D}^{(N1)} & \dots & \mathcal{D}^{(Nj)} & \dots & \mathcal{D}^{(NN)} \end{pmatrix} \cdot \begin{pmatrix} \mathbf{C}^{(1)} \\ \vdots \\ \mathbf{C}^{(j)} \\ \vdots \\ \mathbf{C}^{(N)} \end{pmatrix}$$

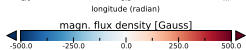
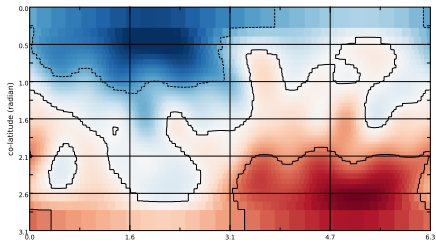
# Example: YY Gem

Magnetic surface maps (Donati et al. 2011)

YY Gem, primary, radial magnetic field



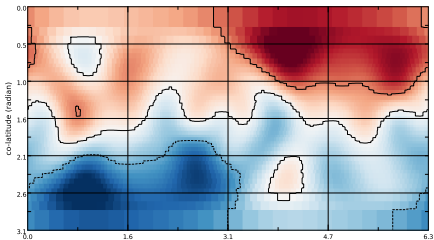
YY Gem, secondary, radial magnetic field



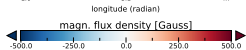
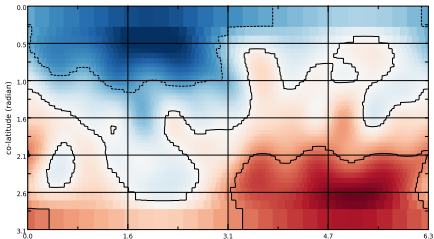
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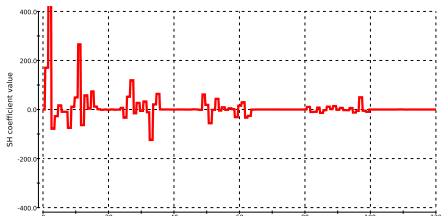
YY Gem, primary, radial magnetic field



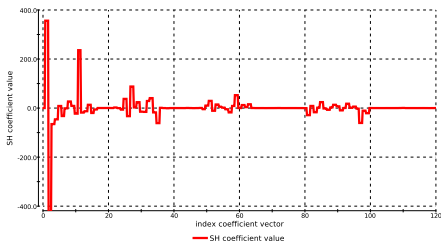
YY Gem, secondary, radial magnetic field



YY Gem, primary, magn. SH coefficient vector



YY Gem, secondary, magn. SH coefficient vector

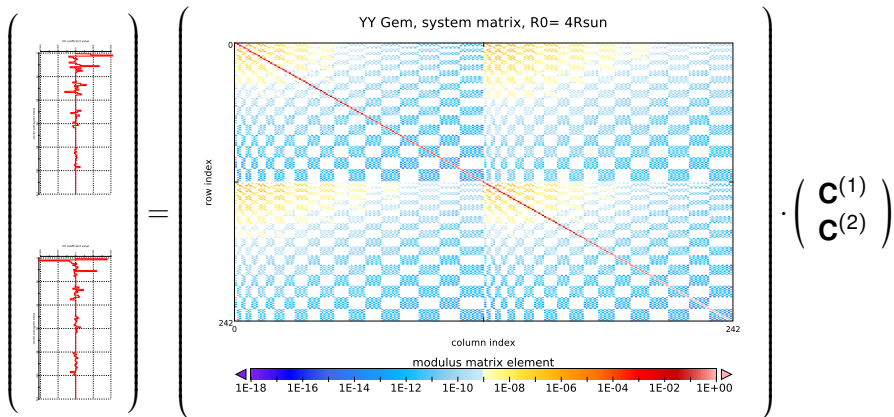


spherical harmonic analyses



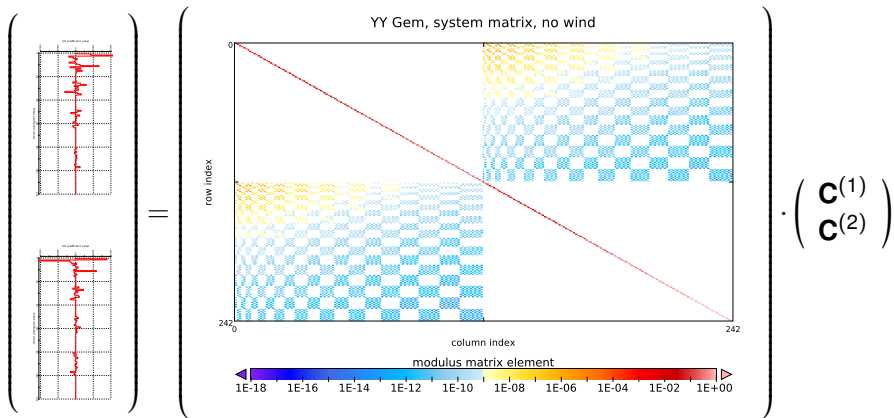
# Example: YY Gem

Linear system of equations ( $R_0 = 4R_\odot$ )



# Example: YY Gem

Linear system of equations ( $R_0 \rightarrow \infty$ )



# Example: YY Gem

Joint magnetosphere – source surface radius  $R_0 \gg d_{\text{YYGem}}$

